

PHYSICS UNIT 3+4 SUMMARY MOTION IN ONE AND TWO DIMENSIONS

x = position (m), t = time (s), v = velocity, a = acc'n
 Δx or s = displacement (change in position) (m), $\Delta x = x_{\text{final}} - x_{\text{initial}}$

If const vel
 $d = vt$

v = velocity (ms^{-1}), $v_{\text{average}} = \frac{\Delta x}{\Delta t}$ a = acceleration (ms^{-2}), $a_{\text{average}} = \frac{\Delta v}{\Delta t}$

$\text{kmh}^{-1} \div 3.6 \rightarrow \text{ms}^{-1}$

Graphs & const acc'n eqns

$x-t$ graph: gradient = velocity **$a-t$ graph:** area = Δv
 $v-t$ graph: gradient = accel'n, area = Δx av. v = area/time

$v = u + at$ $v^2 = u^2 + 2as$ $s = \left(\frac{u+v}{2}\right)t$ $s = ut + \frac{1}{2}at^2$ (*)

(*) or $t = \left(-u \pm \sqrt{u^2 + 2as}\right) \div a$ careful with +, - signs, or find v first, then find t .
 (use $v^2 = \dots$, then $v = u + at$)

When using kinematics eqns, + = initial direction.

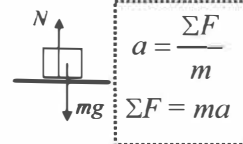
Newton's laws: 1. Inertia. No net force, no change in v ; 2. Net F: $\Sigma F = ma$

3. Action / Reaction forces: $F_{AB} = -F_{BA}$

Gravity: $g = 10 \text{ ms}^{-2}$ or 10 Nkg^{-1} , Weight $W = mg$

Lifts, falling (up = +a): $F_{\text{floor}} - mg = ma$

N = Force of floor on mass = Force of mass on floor

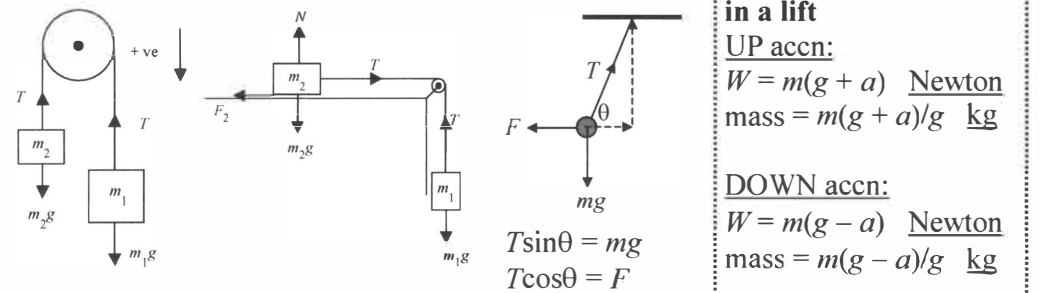
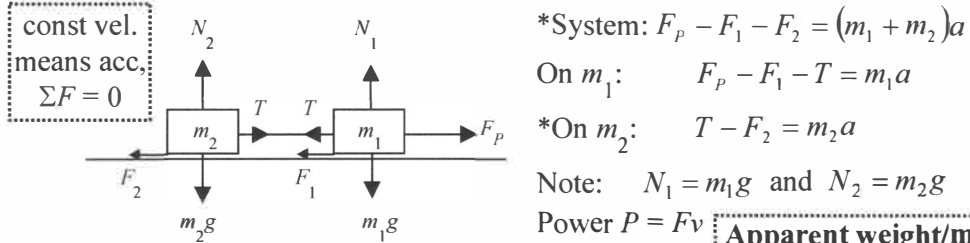


Apparent weight in Newton = F_{floor} , Apparent mass = $F_{\text{floor}} \div g$

Walking: Feet push ground backwards, ground friction pushes you forwd.

Forces and systems. Use $\Sigma F = ma$ If no friction, set $F_1, F_2 = 0$

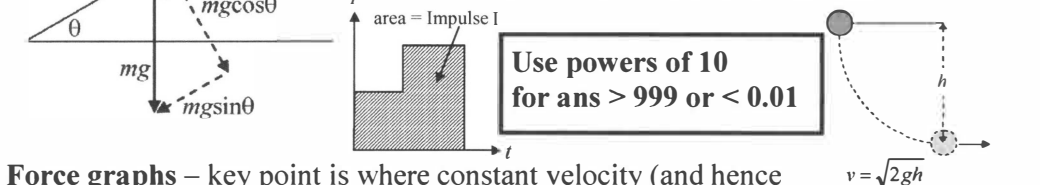
Show ALL forces (including Normal reaction and mg) on each mass. eg:



Inclined plane
 Component of gravity force down plane: $= mg \sin \theta$

acc'n down slope: $a = g \sin \theta$ if no friction
 Normal reaction = $mg \cos \theta$

Friction (if any) acts against motion.



Force graphs – key point is where constant velocity (and hence balanced (equal) driving and resistance forces occur.

Frames of reference / relative velocity $\rightarrow - \downarrow = \rightarrow + \uparrow$ use pythagoras
 Newton's law and motion identical in frame moving at constant velocity (called a inertial frame). velocity of A relative to B: $v_{AB} = v_A - v_B$ (vector add)

Momentum: $P = mv$ (kgms^{-1}), **Impulse:** $I = Ft$ (Ns) or area under $F-t$ graph.
Ext answer: When a car stops in a collision, a certain impulse ($I = Ft$) is required to stop the passenger. An air bag will increase the time t the passenger takes to come to rest. Since I is same, F must decrease, meaning less damage to passenger.

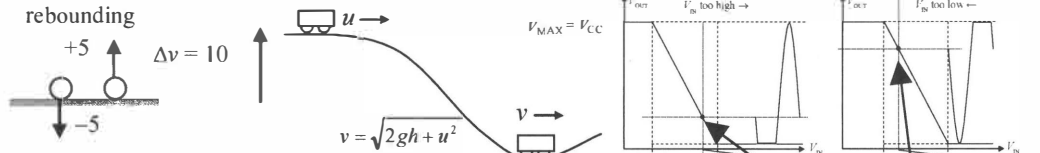
Use numerical values in $Ft = m\Delta v$ or $F = \frac{m\Delta v}{t}$ or $F = \frac{\Delta p}{t}$ to show diff. in F for

two sets of values (e.g. $m = 10 \text{ kg}$, $\Delta v = 20 \text{ ms}^{-1}$, $t = 1 \text{ s}$) & comp. when $t = 0.01$ Use brackets on calculator
 Impulse = Change in momentum $I = \Delta p$ or $Ft = m\Delta v$ or $Ft = m(v-u)$

No external forces \Rightarrow Momentum is always conserved. i.e.:

$\Delta P = 0$, $P_{\text{final}} = P_{\text{initial}}$, mtm before = mtm after, $m_1u_1 + m_2u_2 = m_1v_1 + m_2v_2$

Momentum never lost in elastic or inelastic collisions (but may be transferred to earth).



Kinetic Energy: $KE = \frac{1}{2}mv^2$ (speed $\times 2, 4 \times KE$)

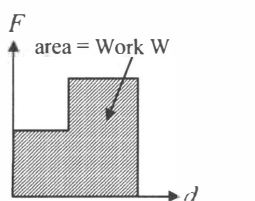
Inelastic collision: $KE_{\text{after}} < KE_{\text{before}}$ Energy lost \rightarrow heat, sound.

e.g. Masses sticking together: $m_1u_1 + m_2u_2 = (m_1 + m_2)v$

Elastic collision: $KE_{\text{after}} = KE_{\text{before}}$

Work $W = Fd$ (if F parallel to d) or $W = (F \cos \theta)d$

Work (Joules) = Area under Force–distance graph.
 If $F \perp d$, no work is done (e.g. in circular motion)



$W = Fd = \Delta KE = \frac{1}{2}mv^2 - \frac{1}{2}mu^2$ or $W = \Delta PE$

$T = \frac{U \sin \theta + \sqrt{(U \sin \theta)^2 + 2gH}}{g}$

Grav'l Pot Energy speed at B: $V = \sqrt{U^2 + 2gH}$

$GPE = mgh$, $mgh = \frac{1}{2}mv^2 - \frac{1}{2}mu^2$

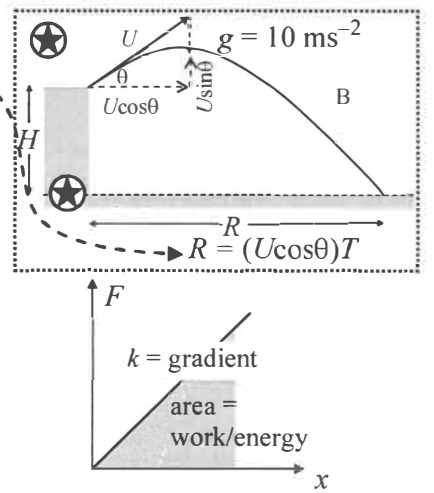
Loss GPE = Gain KE, Gain GPE = Loss KE

Springs: Potential Energy, Hooke's law $F = kx$ (k = stiffness), x = extension or compression only

$SPE = 0.5kx^2$ or Area under $F-x$ graph.

Springs and Hooke's law: $F = -kx$

$SPE = \frac{1}{2}kx^2$



Power $P = W/t$, $P = Fv$ $E = Pt$ Units: Watts or W ($= \text{Js}^{-1}$)

Projectile motion Accel'n occurs only vertically. Horiz. velocity constant.

Find components of initial velocity U $d_{\text{Horiz}} = v_{\text{Horiz}} \times t$, $d_H = U \cos \theta \times t$

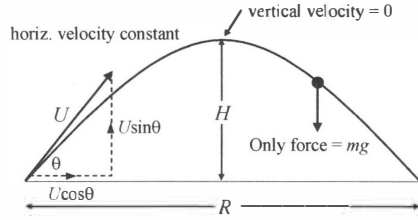
(Vertically: $u = U \sin \theta$, $a = -g = -10$, Horizontally: $u = U \cos \theta$, $a = 0$)

Use Kinematics eqns (a) Vert., and (b) Horiz.

May also use: $mgh = \frac{1}{2}mv^2 - \frac{1}{2}mu^2$

tonne
t = 1000 kg

(i) Half parabola (from a level, up, then down to same level).



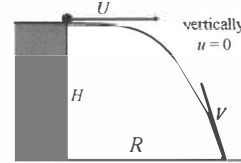
Range: $R = (U \cos \theta) T$

$R = \frac{U^2 \sin(2\theta)}{g}$

$H = \frac{(U \sin \theta)^2}{(2g)}$

Time to max. height: $t = \frac{U \sin \theta}{g}$, Total time ($\uparrow + \downarrow$): $T = \frac{2U \sin \theta}{g}$

(ii) Horizontal projection from cliff of height h .



$R = UT$, $R = U \sqrt{\frac{2H}{g}}$, $V = \sqrt{U^2 + 2gH}$

Time of flight $T = \sqrt{\frac{2H}{g}}$, $H = \frac{T^2 g}{2}$

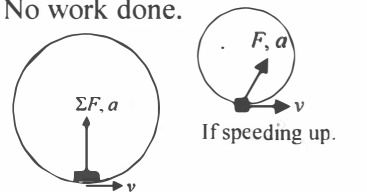
(iii) Oblique proj'n from cliff of height H . Use kin eqns, + = initial direction, or

Circular motion – uniform (constant speed) Velocity: $v = \frac{2\pi R}{T}$, $T = \sqrt{\frac{m4\pi^2 R}{F}}$

Sum of forces is towards centre. Velocity is tangential. No work done.

$F = \frac{mv^2}{R}$ $F = \frac{m4\pi^2 R}{T^2}$ $F = \frac{m2\pi v}{T}$ $F = m4\pi^2 R f^2$

$a = \frac{v^2}{R}$ $a = \frac{4\pi^2 R}{T^2}$ $a = \frac{2\pi v}{T}$ $a = 4\pi^2 R f^2$



Gravity and Newton's law of univ'l gravit'n (see also circ motion above)

$G = 6.67 \times 10^{-11}$, M = central body, m = orbiting obj't

$M_{\text{earth}} = 5.98 \times 10^{24} \text{ kg}$, $R_{\text{earth}} = 6.37 \times 10^6 \text{ m}$, $M_{\text{sun}} = 2 \times 10^{30} \text{ kg}$

1 day = 86400 s, 1 year = $3.15 \times 10^7 \text{ s}$ Weight W = Force F , field (g) = acc'n a

$v = \frac{2\pi R}{T}$ $T = \frac{2\pi R}{v}$ $R = \frac{vT}{2\pi}$ Add height above planet to radius to find orbital Radius if necessary.

$F = \frac{GMm}{R^2}$ grav. field: $g = \frac{GM}{R^2}$, $a = \frac{GM}{R^2}$, $g = \frac{v^2}{R}$, $a = \frac{v^2}{R}$

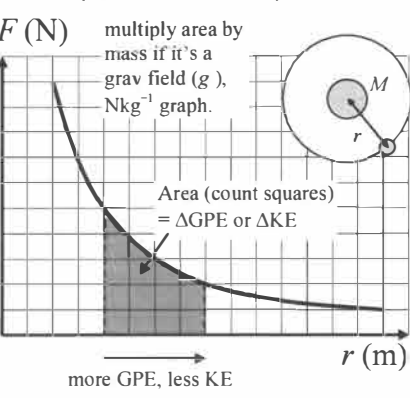
$F = \frac{mv^2}{R} = \frac{GMm}{R^2} \Rightarrow v = \sqrt{\frac{GM}{R}}$, $R = \frac{GM}{v^2}$, $M = \frac{v^2 R}{G}$, $KE = \frac{1}{2}mgR$

$F = \frac{m4\pi^2 R}{T^2} = \frac{GMm}{R^2} \Rightarrow \frac{R^3}{T^2} = \frac{GM}{4\pi^2} = \text{constant for same central body } M$

$R = \sqrt[3]{\frac{GMT^2}{(4\pi^2)}}$ $T = \sqrt{\frac{4\pi^2 R^3}{GM}}$ $T = \sqrt{\frac{4\pi^2 R}{g}}$ $M = \frac{4\pi^2 R^3}{(GT^2)}$ $v = \sqrt{\frac{2\pi GM}{T}}$

Ratio questions: Two masses m_1, m_2 orbiting central body M .

$\frac{T_1}{T_2} = \sqrt{\left(\frac{R_1}{R_2}\right)^3}$, $\frac{R_1}{R_2} = \sqrt[3]{\left(\frac{T_1}{T_2}\right)^2}$, $\frac{a_1}{a_2} = \frac{g_1}{g_2} = \left(\frac{R_2}{R_1}\right)^2$, $\frac{v_1}{v_2} = \sqrt{\frac{R_2}{R_1}}$, $\frac{F_1}{F_2} = \frac{m_1}{m_2} \left(\frac{R_2}{R_1}\right)^2$



$F = \frac{GMm}{R^2}$, $v = \frac{2\pi R}{T}$

$KE = \frac{1}{2}mv^2 = \frac{1}{2}mgR$

k = 10³
M = 10⁶
G = 10⁹

*If using g field graph (Nkg^{-1}), multiply by mass to get ΔPE or ΔKE or Work.

Geostationary (geosynchronous) orbit:
 Period T of satellite = period (length of "day" in sec) of rotation of planet. $T_{\text{earth}} = 86400 \text{ s}$

Unit 4 Properties of mechanical waves

- $T = \frac{1}{f}$

- $v = f\lambda = \frac{\lambda}{T}$

- Doppler effect: $f = \frac{v_{\text{sound}}}{\left(\frac{v_{\text{sound}} \pm v}{f_0}\right)}$ P.19

= $\frac{v_{\text{sound}} f_0}{v_{\text{sound}} \pm v}$

- harmonics

fixed at both ends:

$\lambda = \frac{2L}{n} \rightarrow f = \frac{nv}{2L}$

$n = 1, 2, 3, 4$

- fixed at one end

odd $\lambda = \frac{4L}{n} \rightarrow f = \frac{nv}{4L}$ ($n = 1, 3, 5$)

n : number of harmonics (only odd-numbered harmonics)

$\lambda = \frac{4L}{2n-1}$ $n = 1, 2, 3, 4, \dots$

n : the next harmonic in the sequence

not harmonic number

- refractive index

$n = \frac{c}{v}$

- Snell's law $n_1 \sin \theta_1 = n_2 \sin \theta_2$

- internal reflection.

critical angle - $n_1 \sin \theta_c = n_2 \sin 90^\circ$ ($n_1 > n_2$)

- constructive interference $Pd = n\lambda$ ($n = 0, 1, 2, \dots$)

- destructive interference $Pd = n\lambda$ ($n = 0.5, 1.5, 2.5, \dots$)

- fringe separation

$\Delta x = \frac{\lambda}{\theta}$

Planck's equation

- $E = hf \Rightarrow f = \frac{c}{\lambda}$

- $E = \frac{hc}{\lambda}$

- $E_{\text{max}} = hf - \phi$

$\phi = hf_0$

\rightarrow stopping voltage $V \Rightarrow E = eV_{\text{stopping}}$
the charge on an electron

$eV_{\text{stopping}} = E_k = \frac{1}{2} m v^2$ speed

- $\lambda = \frac{h}{p} = \frac{h}{mv}$

- $\Delta E = E_n - E_m = hf = \frac{hc}{\lambda}$

(n, m Energy level)

- De Broglie wave

$mvr = n \frac{h}{2\pi}$

$C = n\lambda$

- uncertainty.

$\Delta x \Delta p \geq \frac{h}{4\pi}$

$\Delta E \cdot \Delta t \geq \frac{h}{4\pi}$

(1 second)

- number of photon per second \uparrow

$n = \frac{\text{Power of light} \cdot \Delta t}{\text{Energy of 1 photon}}$

Energy of 1 photon

- $F = \frac{\Delta p}{\Delta t} \rightarrow \text{momentum}$

= $\frac{P}{c} \rightarrow \text{power}$

- fringe separation

if it is reflected $F = \frac{2P}{c}$